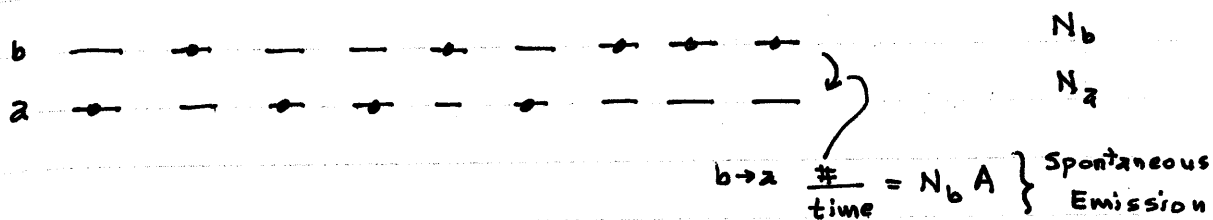


Spontaneous Emission

N_b atoms where e^- is in the "higher energy" state



$$b \rightarrow a \frac{\#}{\text{time}} = N_b A$$

Stimulated Emission $R_{b \rightarrow a} = \frac{\pi}{3\epsilon_0 \hbar^2} |\rho|^2 \rho(\omega_0)$ } incoherent unpolarized all directions

$$b \rightarrow a \frac{\#}{\text{time}} = N_b B_{ba} \rho(\omega_0)$$
 } Stimulated Emission

$$B_{ba} = \frac{\pi}{3\epsilon_0 \hbar^2} |\rho|^2$$

Absorption Rate $a \rightarrow b \frac{\#}{\text{time}} = N_a B_{ab} \rho(\omega_0)$

In an ambient field of e.m. radiation

$$\frac{dN_b}{dt} = -N_b A - N_b B_{ba} \rho(\omega_0) + N_a B_{ab} \rho(\omega_0)$$

If the atoms are in thermal equilibrium with the ambient field, then

$$0 = -N_b A - N_b B_{ba} \rho(\omega_0) + N_a B_{ab} \rho(\omega_0)$$

$$\rho(\omega_0) = \frac{A}{\frac{N_a}{N_b} B_{ab} - B_{ba}}$$

Compare this to Planck Black Body

Spontaneous Emission

First of all, from statistical mechanics, we have the Boltzmann factor

$$\frac{N_a}{N_b} = \frac{e^{-E_a/kT}}{e^{-E_b/kT}} = e^{(E_b - E_a)/kT} = e^{\hbar\omega_0/kT}$$

So, $\rho(\omega_0)$ becomes:
$$\rho(\omega_0) = \frac{A}{e^{\hbar\omega_0/kT} B_{ab} - B_{ba}}$$

Compare this to Black Body Radiation:
$$\rho(\omega) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\hbar\omega/kT} - 1}$$

Comparing these two equations we see that

$$B_{ab} = B_{ba}$$

and

spont. emission rate

$$A = \frac{\omega_0^3 \hbar}{\pi^2 c^3} B_{ba}$$

$$A = \frac{\hbar \omega_0^3}{\pi^2 c^3} \left(\frac{\pi |p|^2}{3\epsilon_0 \hbar^2} \right)$$

$$A = \frac{\omega^3}{3\pi\epsilon_0 \hbar c^3} |p|^2$$

Spontaneous Emission Rate.

Homework Problems: 9.8 9.10 9.11

9.11 You need to calculate $|p|^2$ where $p = \langle \text{final} | e^{\vec{r}} | \text{initial} \rangle$
and $\vec{r} = x\hat{i} + y\hat{j} + z\hat{k}$ $p = e \langle \text{final} | x\hat{i} + y\hat{j} + z\hat{k} | \text{initial} \rangle$

z matrix element $p \rightarrow 0$ except for $\langle 100 | z | 210 \rangle = \frac{2^8}{3^5 \sqrt{2}}$
 $|100\rangle$ $|200\rangle$ and $|210\rangle$ are even in (x, y)

Spontaneous Emission

Problem 9.11 continued.

So the only ^{other} non-zero matrix elements are:

$$\langle 100 | x | 21 \pm 1 \rangle \quad \text{and} \quad \langle 100 | y | 21 \pm 1 \rangle$$

$$\omega = \frac{E_2 - E_1}{\hbar} = -\frac{3E_1}{4\hbar}$$

Calculate A

The lifetime is $\tau = \frac{1}{A}$